

Next Generation of NIKA2 Pixels

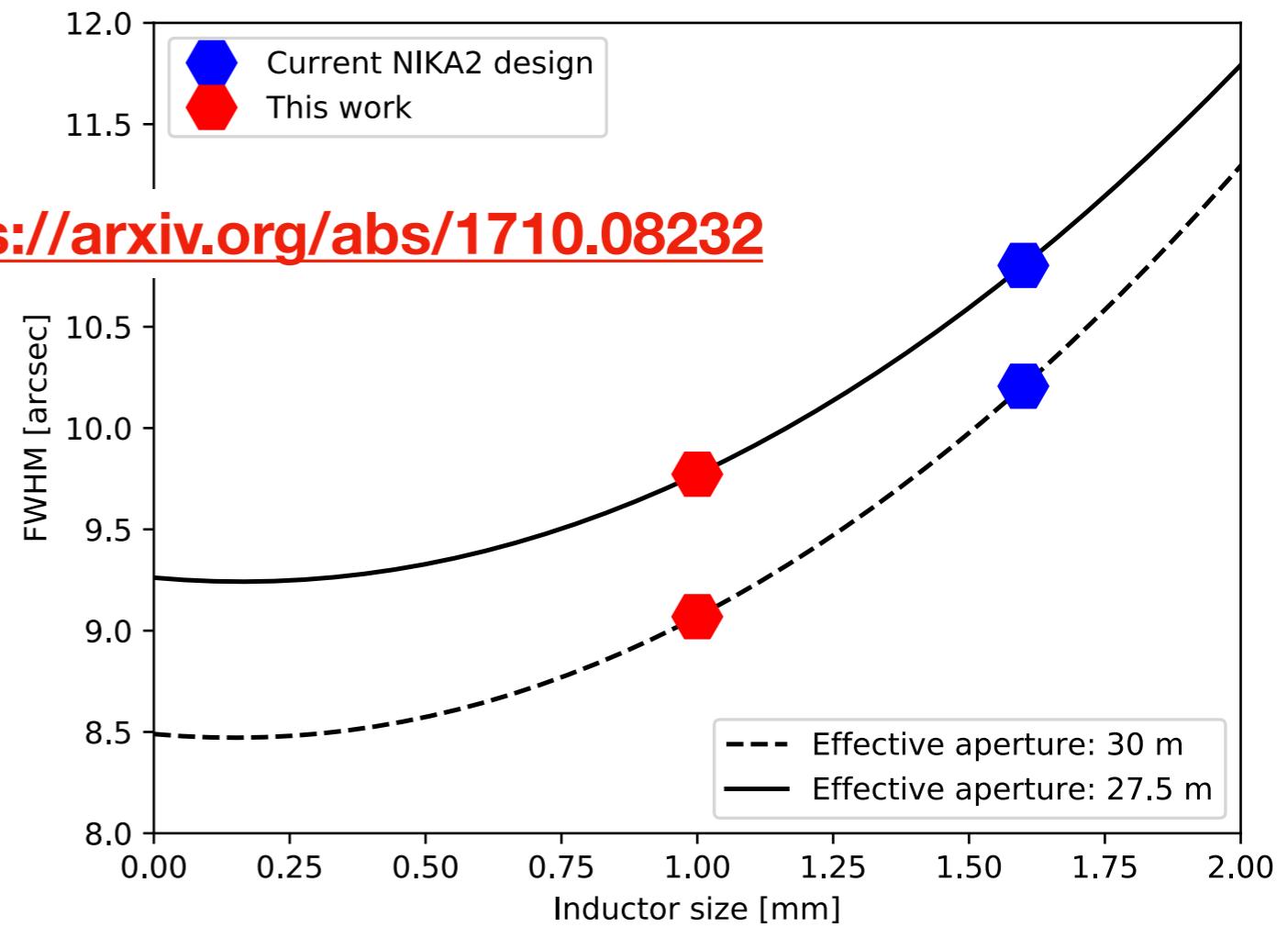


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Calvo, J. Goupy and E. Driessen



Future Development for NIKA2

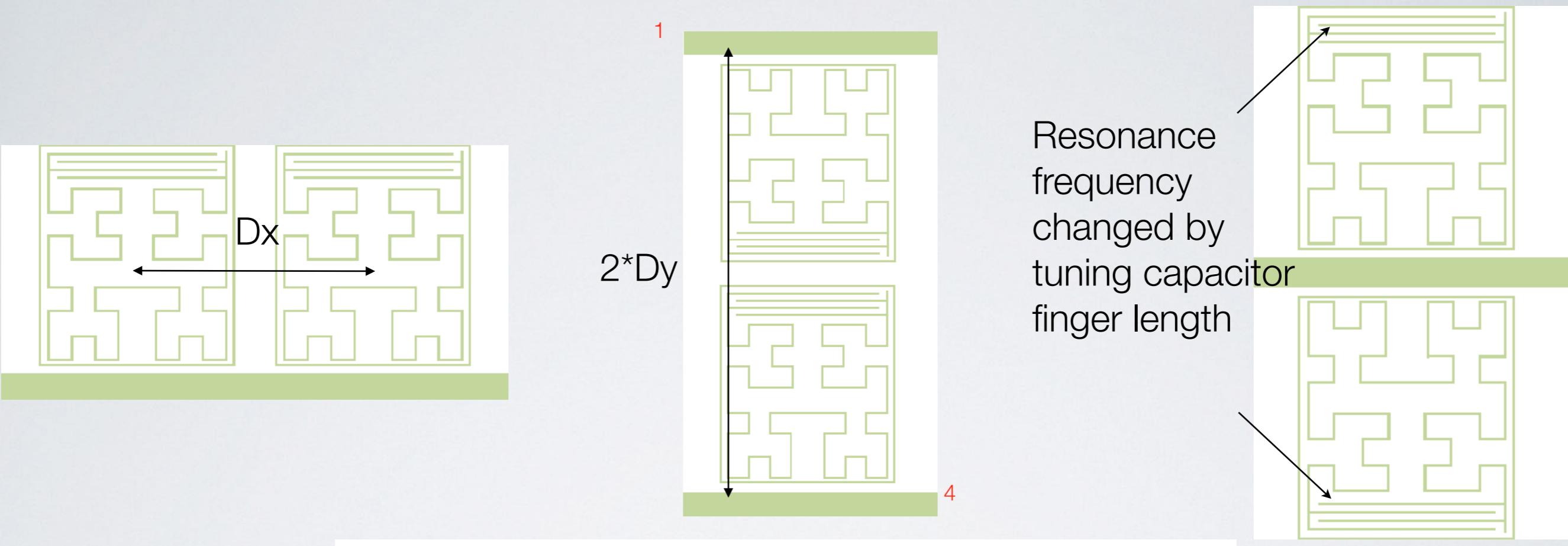
- Higher angular resolution:
10.8" -> 9.8" @1mm
Updated results: <https://arxiv.org/abs/1710.08232>
- Larger band coverage:
230-290 GHz -> 200-310 GHz
- More sensitive:
current 1mm array has 25
 $\text{mJy}\cdot\text{s}^{1/2}$



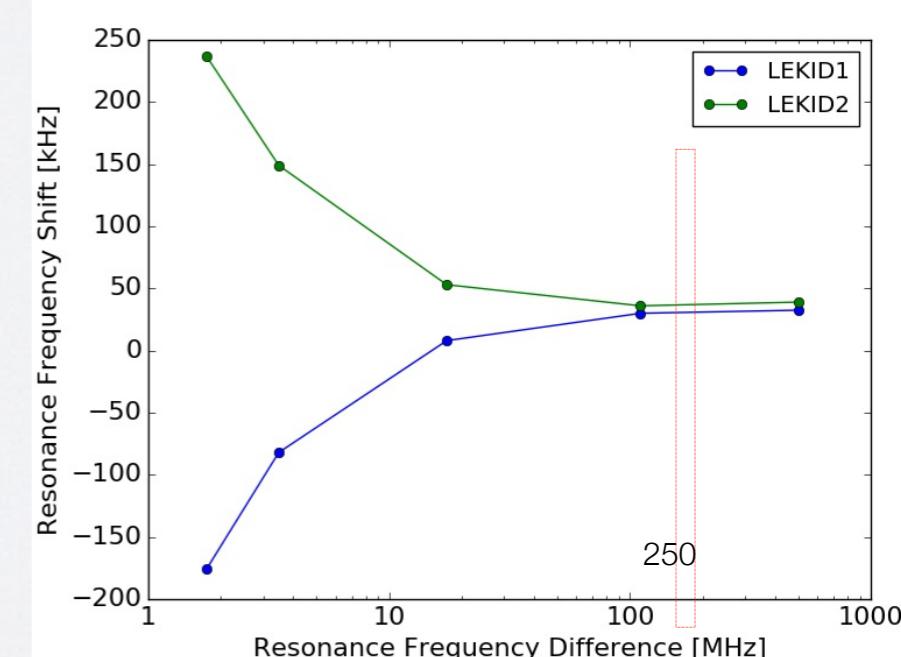
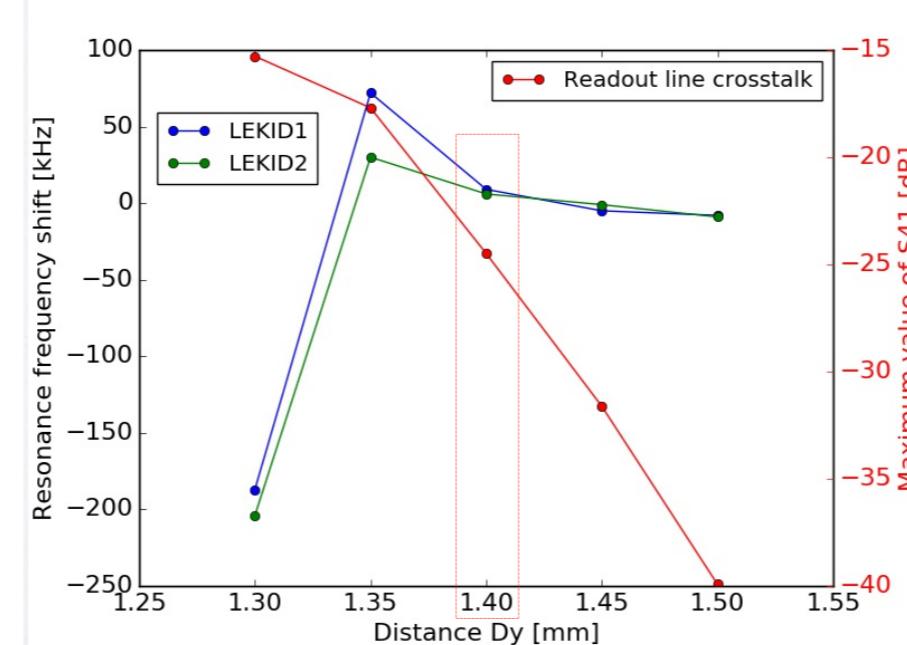
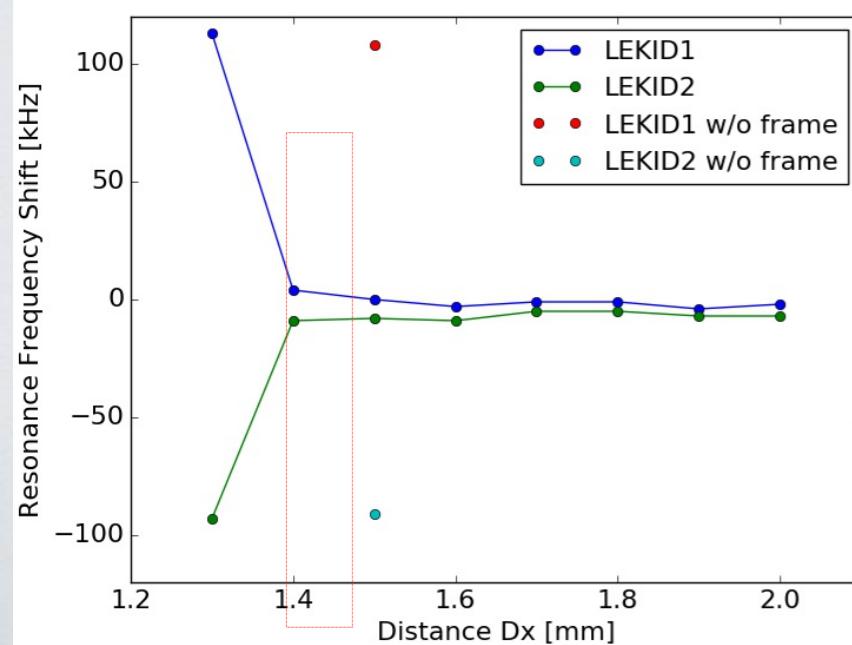
Future Development for NIKA2

- Higher angular resolution:
10.8" -> 9.8" @1mm **Prototype High Angular Resolution LEKIDs for NIKA2**
Updated results: <https://arxiv.org/abs/1710.08232>
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 $\text{mJy}\cdot\text{s}^{1/2}$ **Optical Response Analysis of LEKIDs Array**

Prototype high resolution LEKIDs

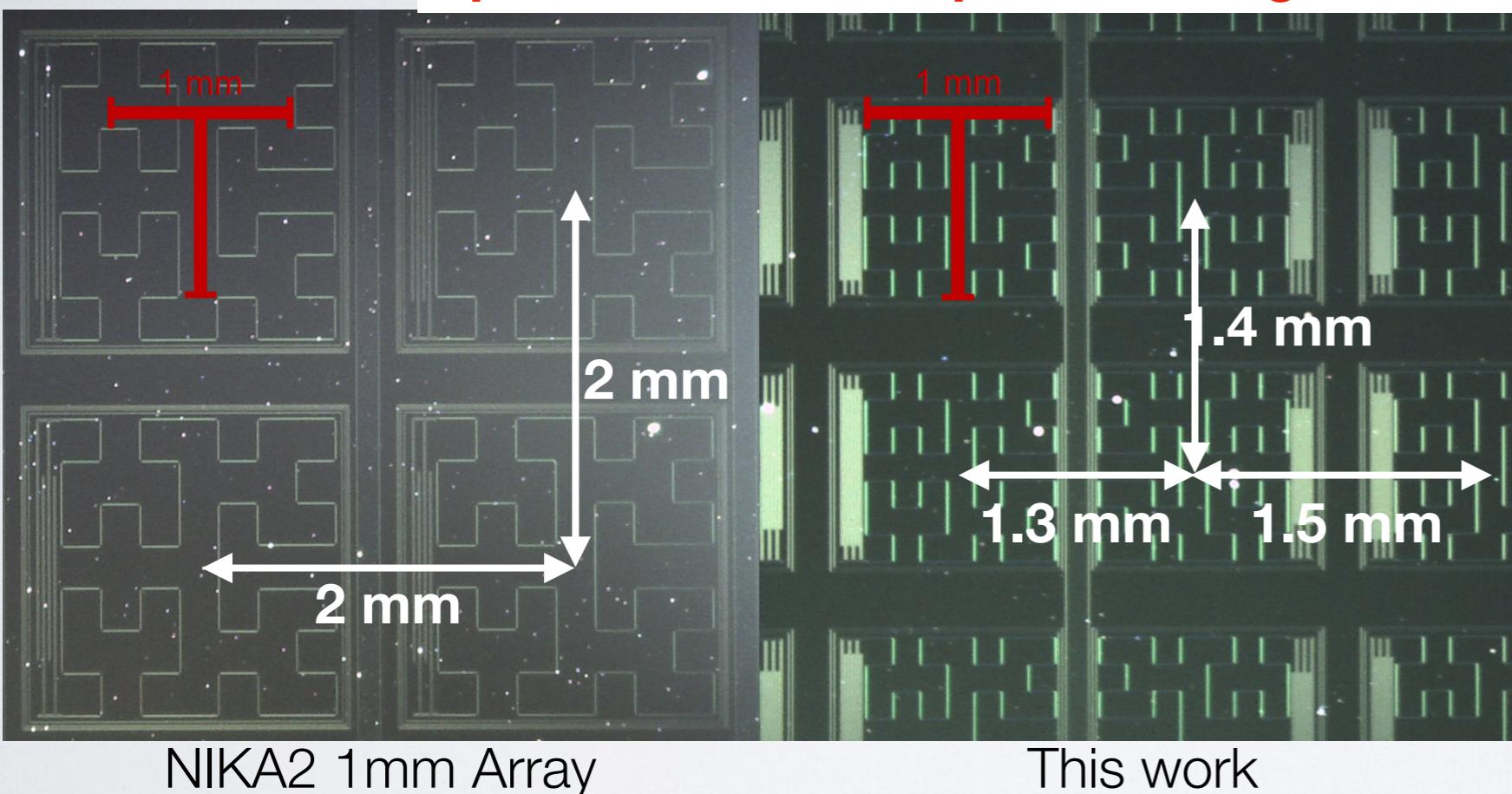


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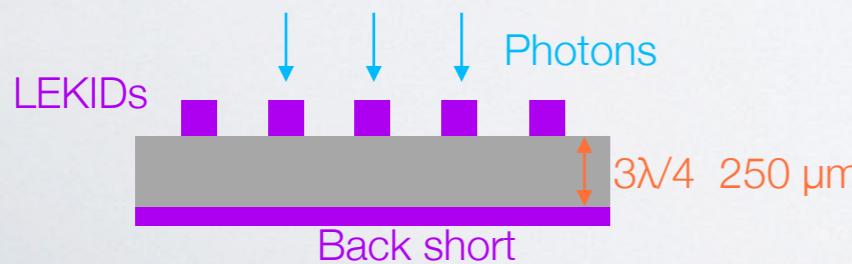


Prototype high resolution LEKIDs

Updated results: <https://arxiv.org/abs/1710.08232>



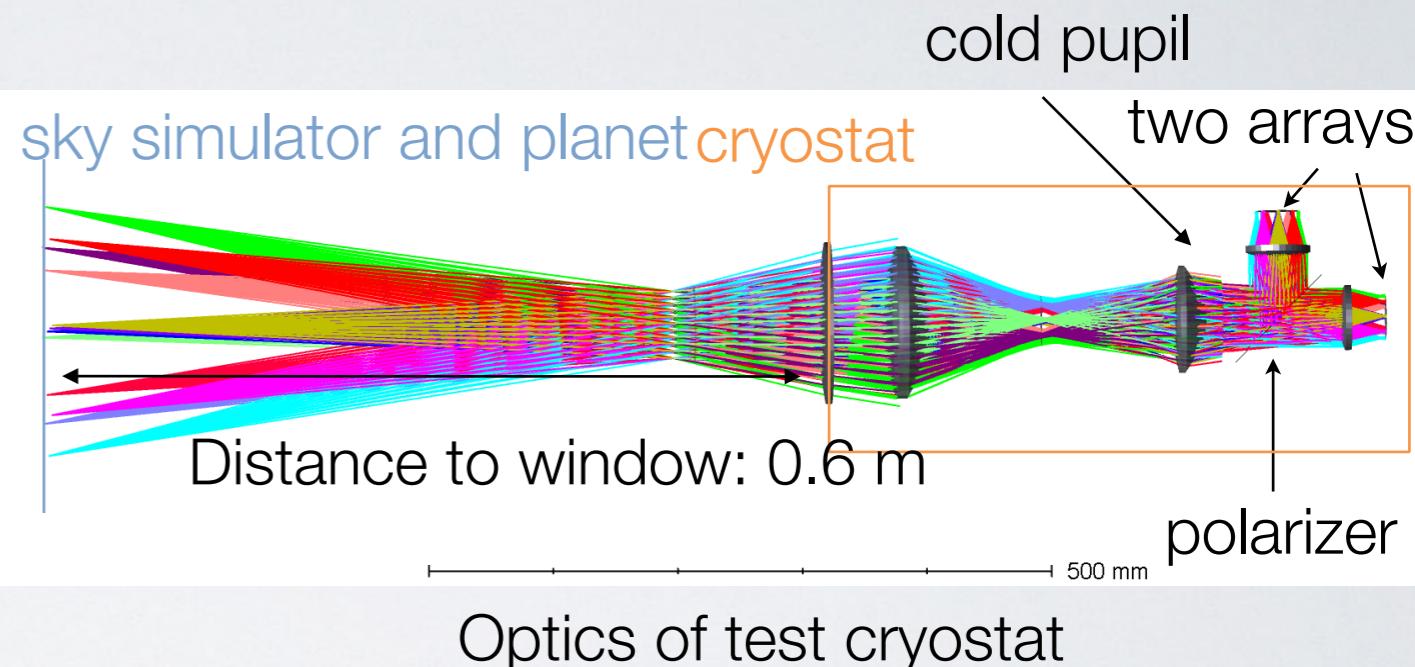
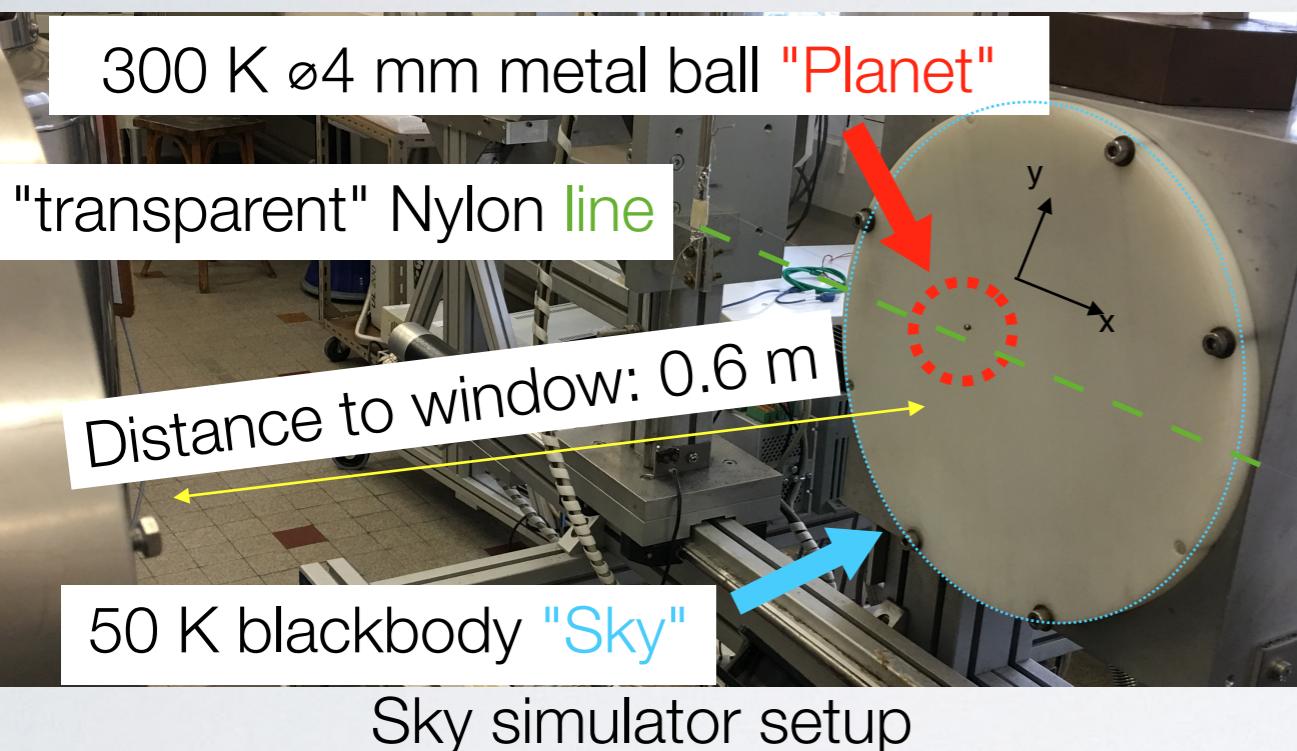
| | NIKA2 1 mm Array | This work |
|---------------------|-------------------------|-------------------------|
| Inductor size | 1.5*1.6 mm ² | 1*1 mm ² |
| Pixel size | 2*2 mm ² | 1.4*1.4 mm ² |
| Resonance frequency | 1.9-2.4 GHz | 2.1-2.6 GHz |
| Inductor width | 4 μ m | 2.5 μ m |
| Resolution [FWHM] | 10.80" | 9.77" |



- 20 nm Al
- 200 nm Al backshort

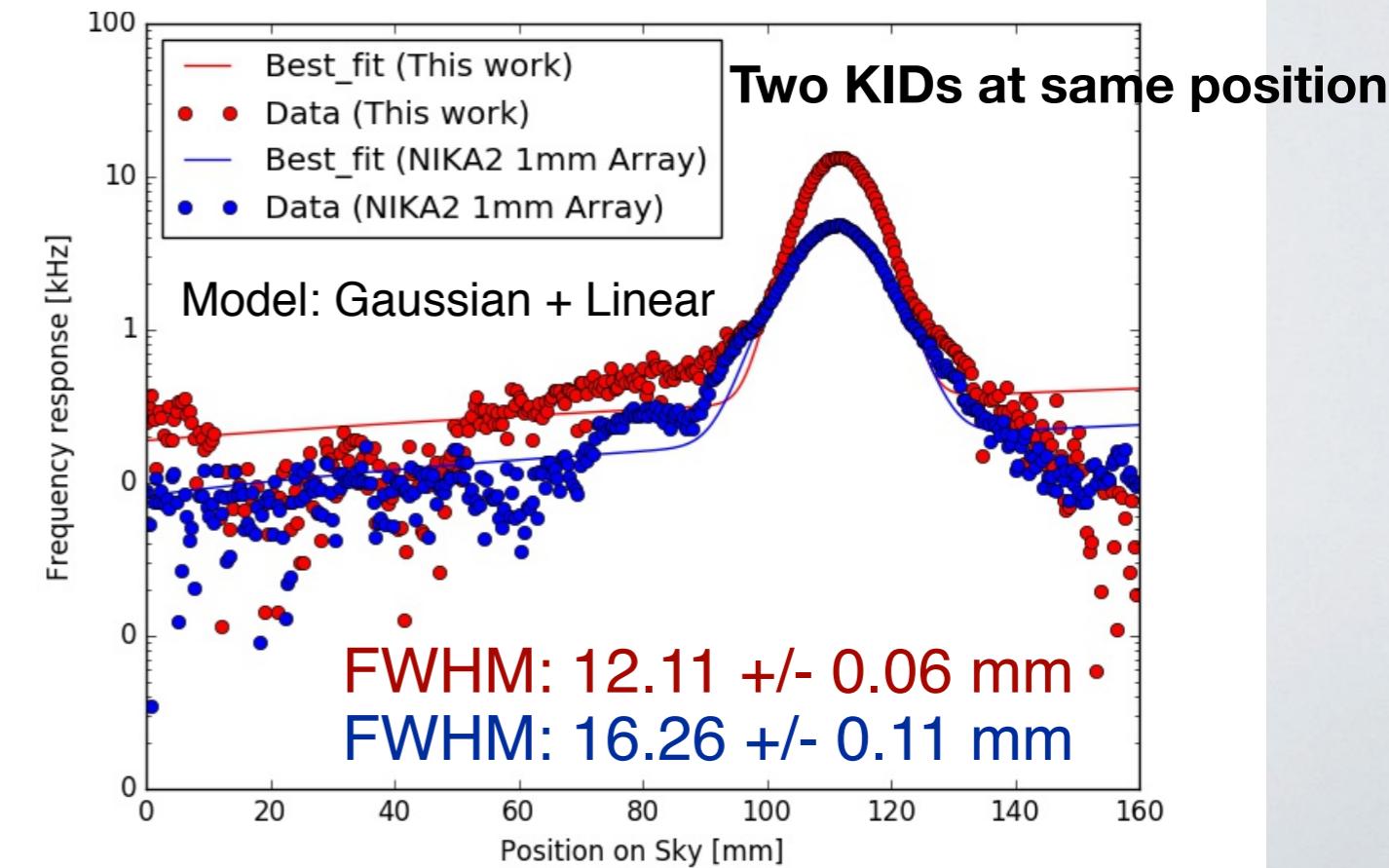
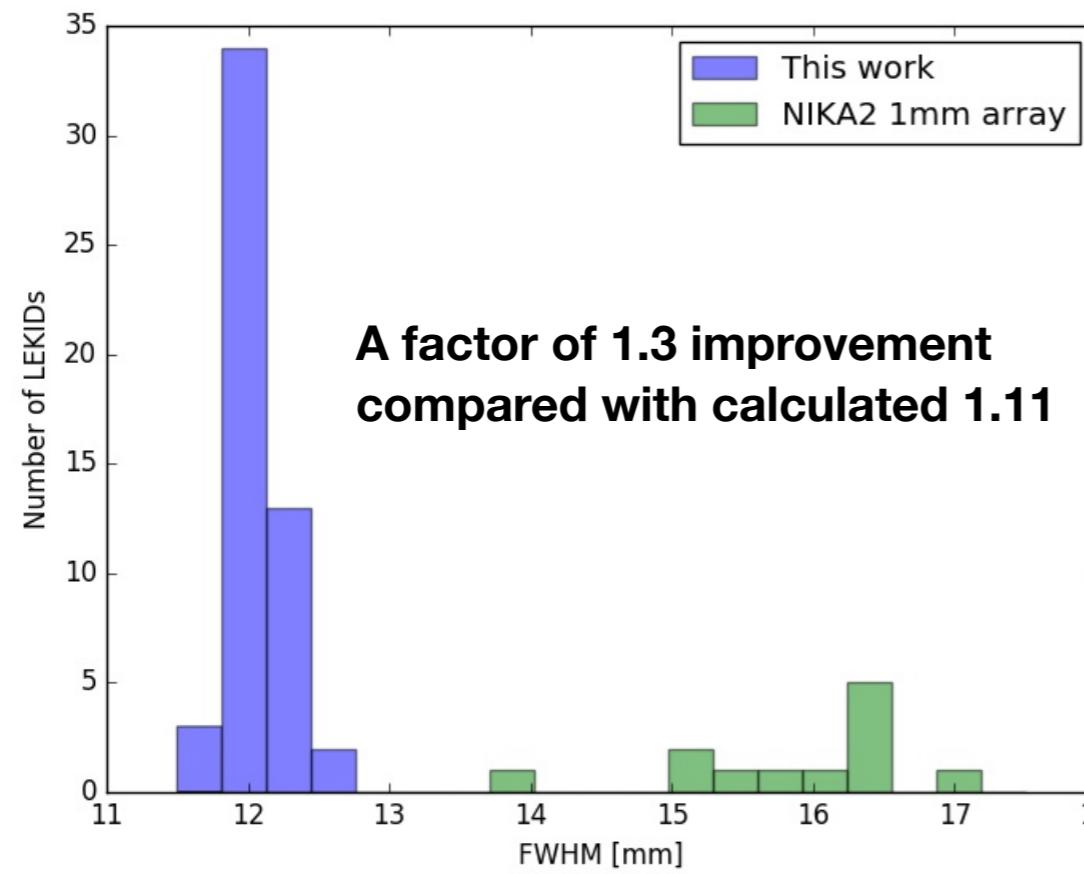
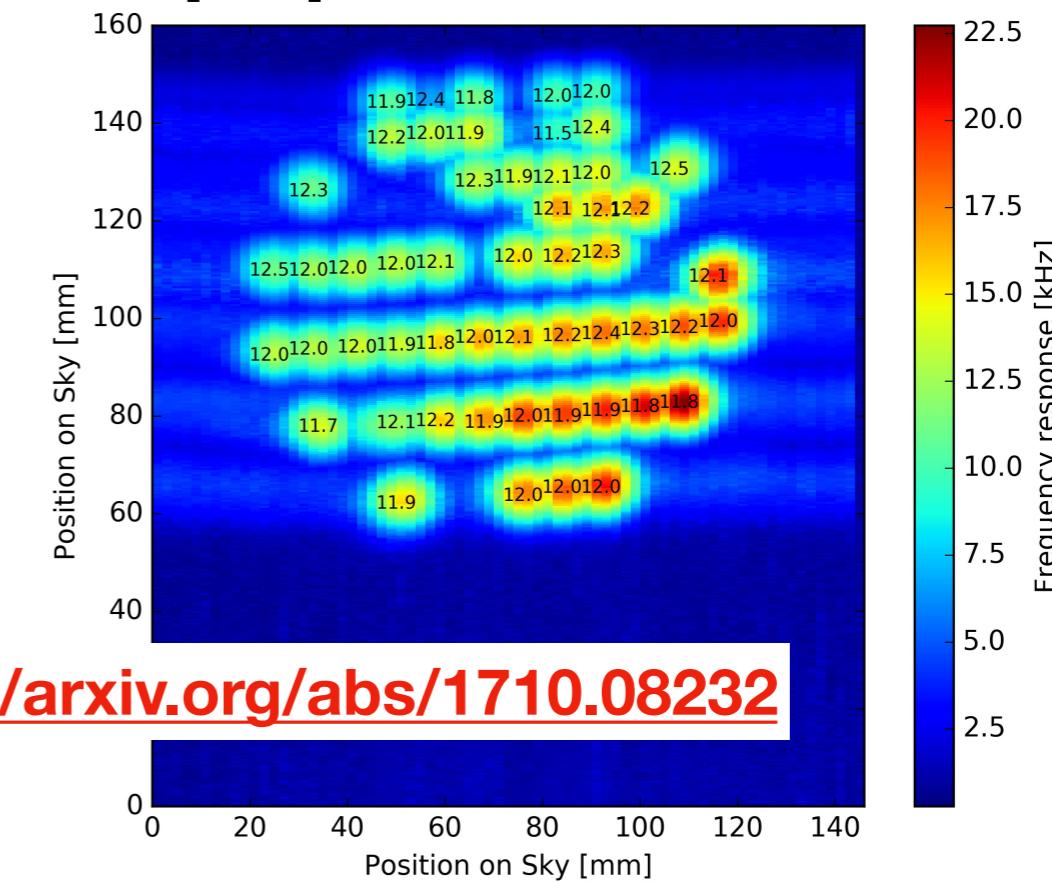
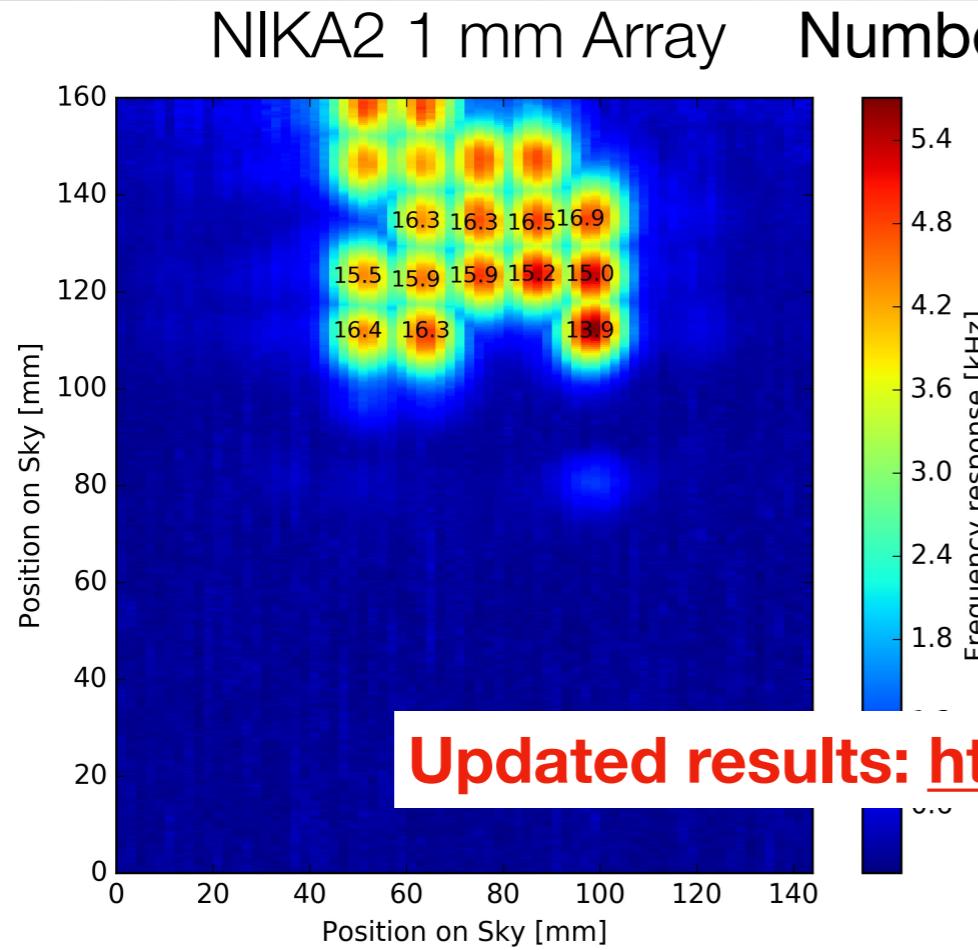
Measurement optics

Updated results: <https://arxiv.org/abs/1710.08232>



- The optics and sky simulator optically behave the same as the NIKA2 at the 30-m telescope:
the planet image on focal plane \leftrightarrow point spread function on telescope
- This optics has a F#1.48 and a magnification factor of 0.38

Measurement result



Optical response simulation of LEKIDs array

Optical response of LEKID

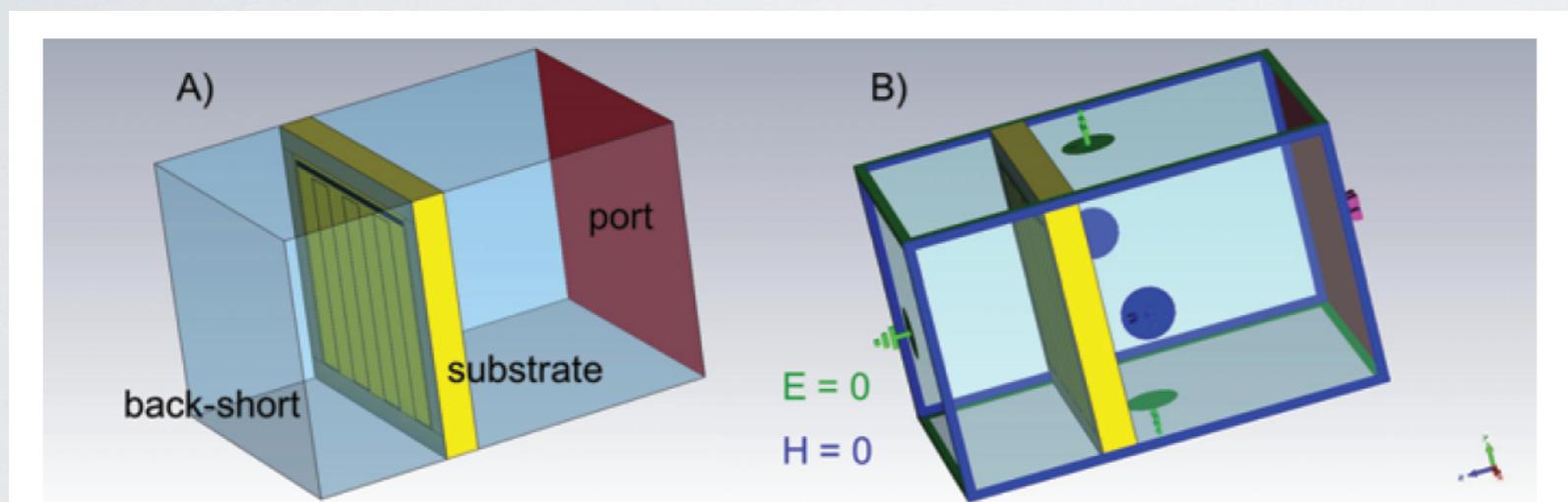
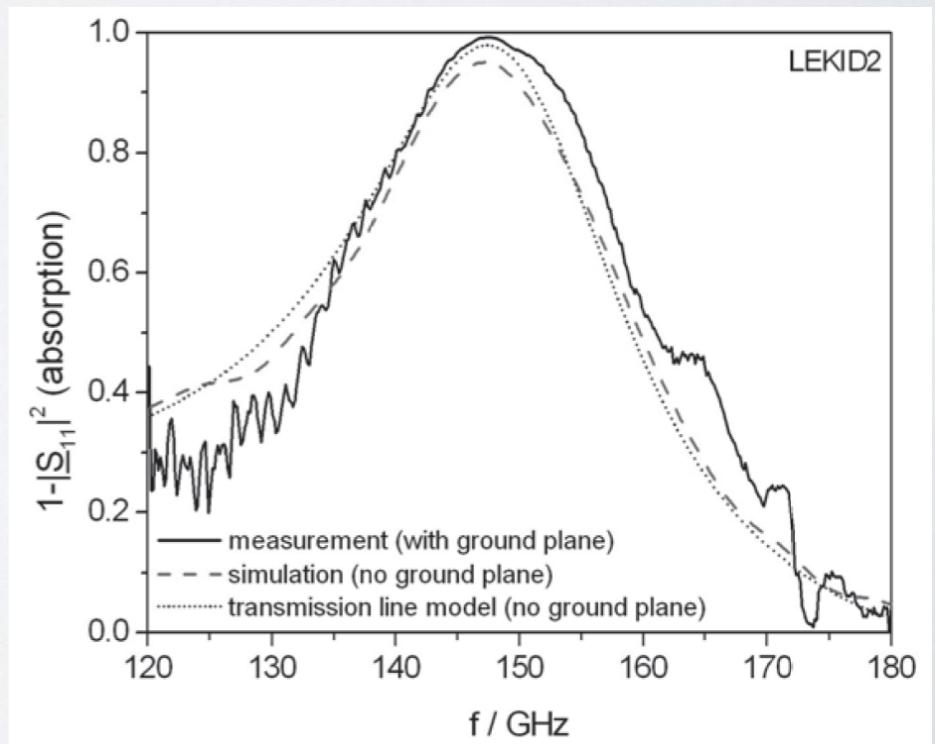
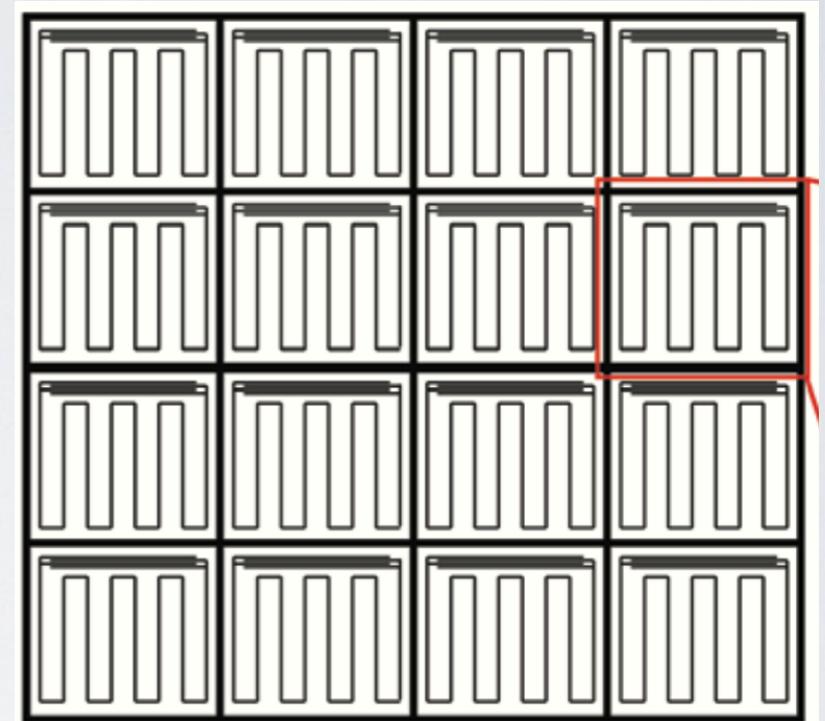
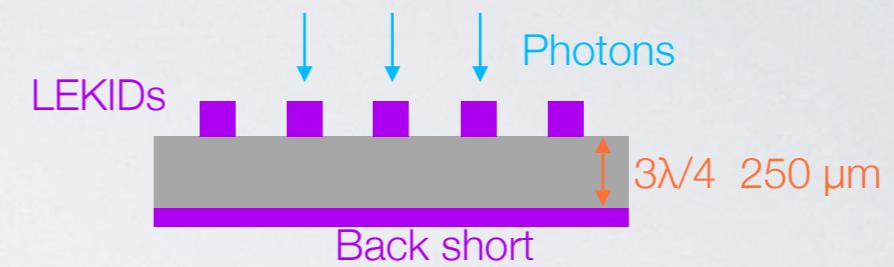
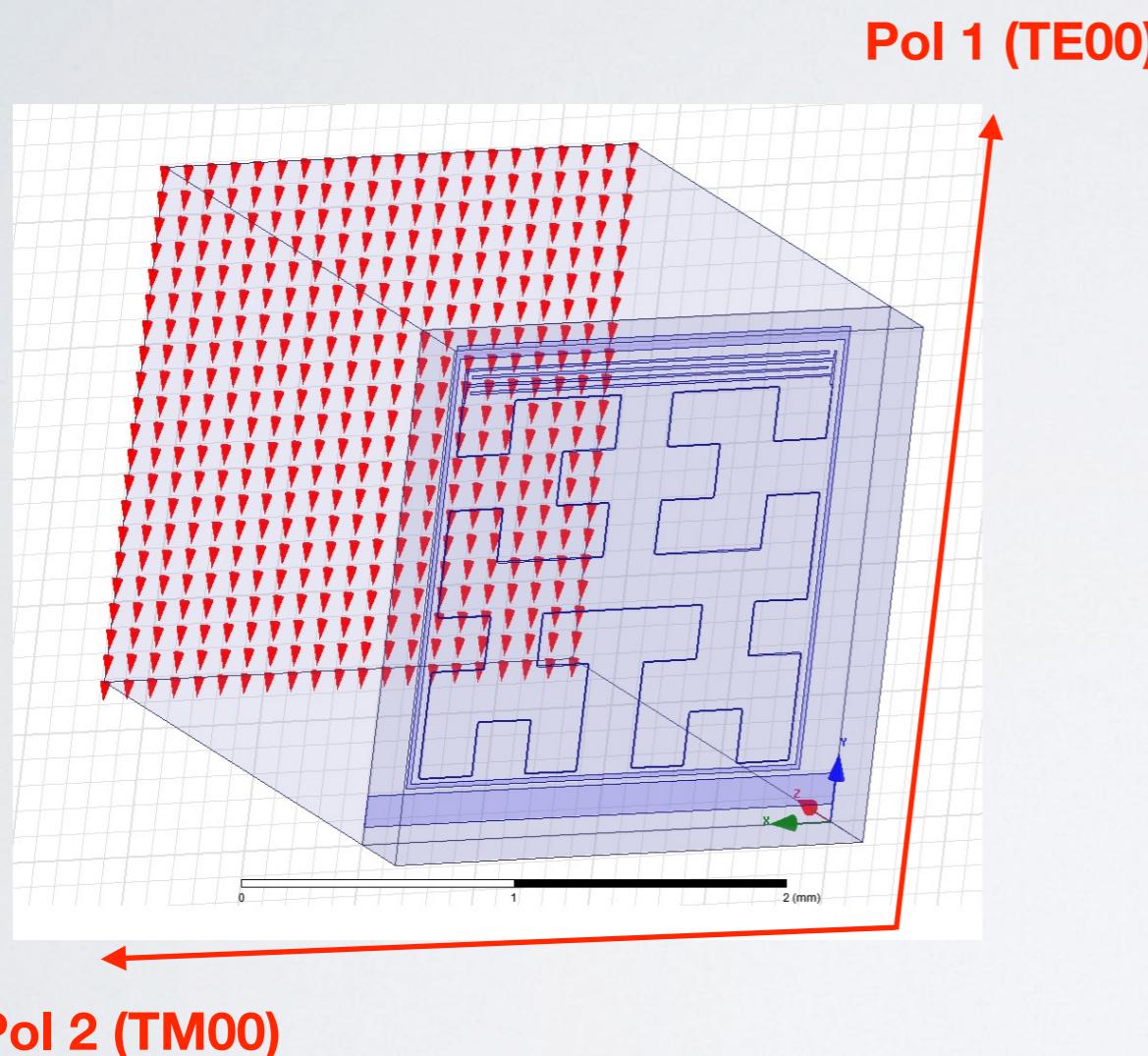


Fig. 7.21.: A) Schematic of the CST model to simulate the optical coupling of LEKIDs. B) The model is considered as a waveguide defined by the boundary conditions $H = 0$ (blue) and $E = 0$ (green). The back-short is modeled as an electrically shorted wall with $E = 0$.

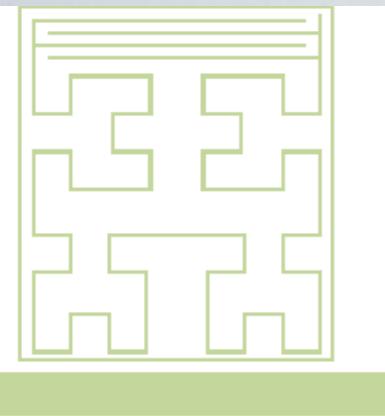
Rosch, M. J., et al. (2013). *IEEE Transactions on Antennas and Propagation*,



Simulation setup

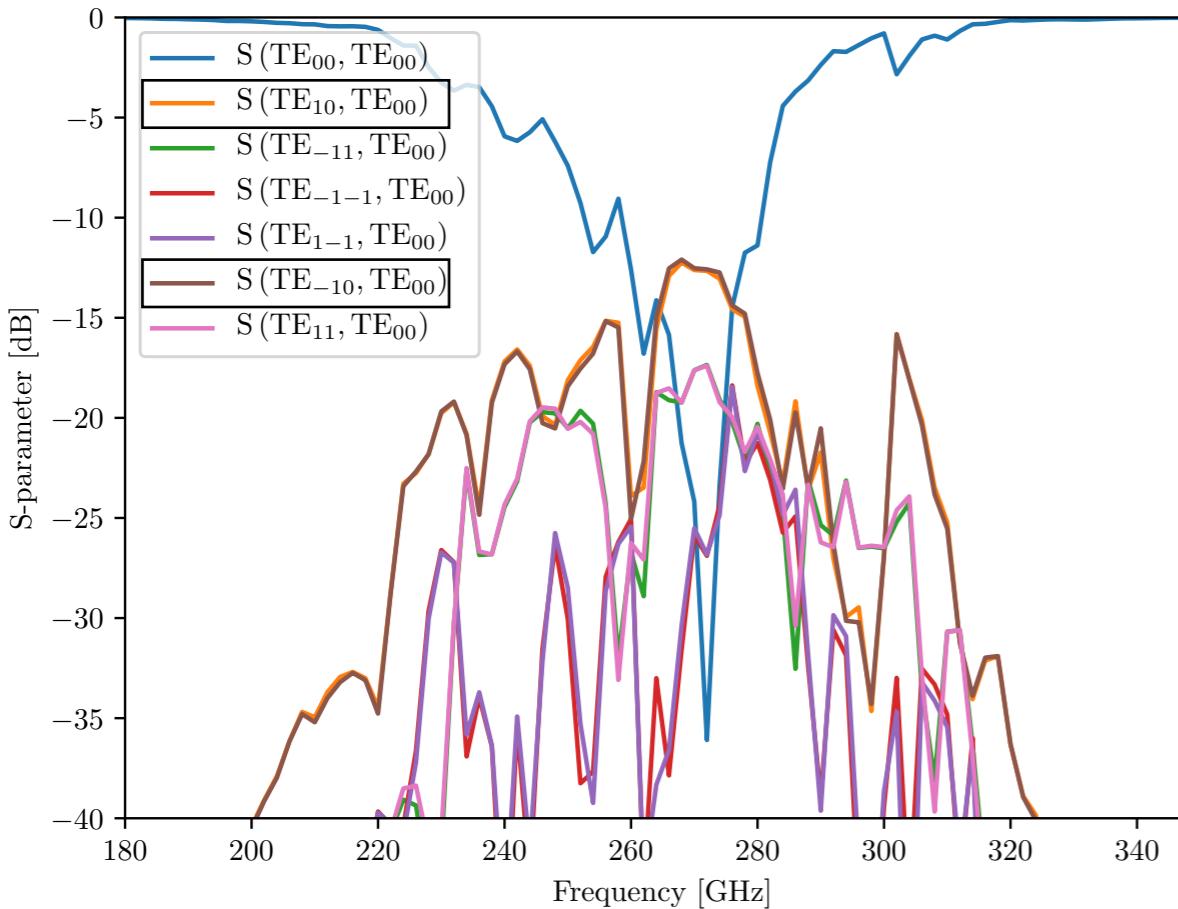


- Floquet port with Periodic boundaries
- 18 modes for 150 - 350 GHz
- Two polarizations incident
- 1.6 Ohm/square for 20 nm Al
- 250 μm Si substrate

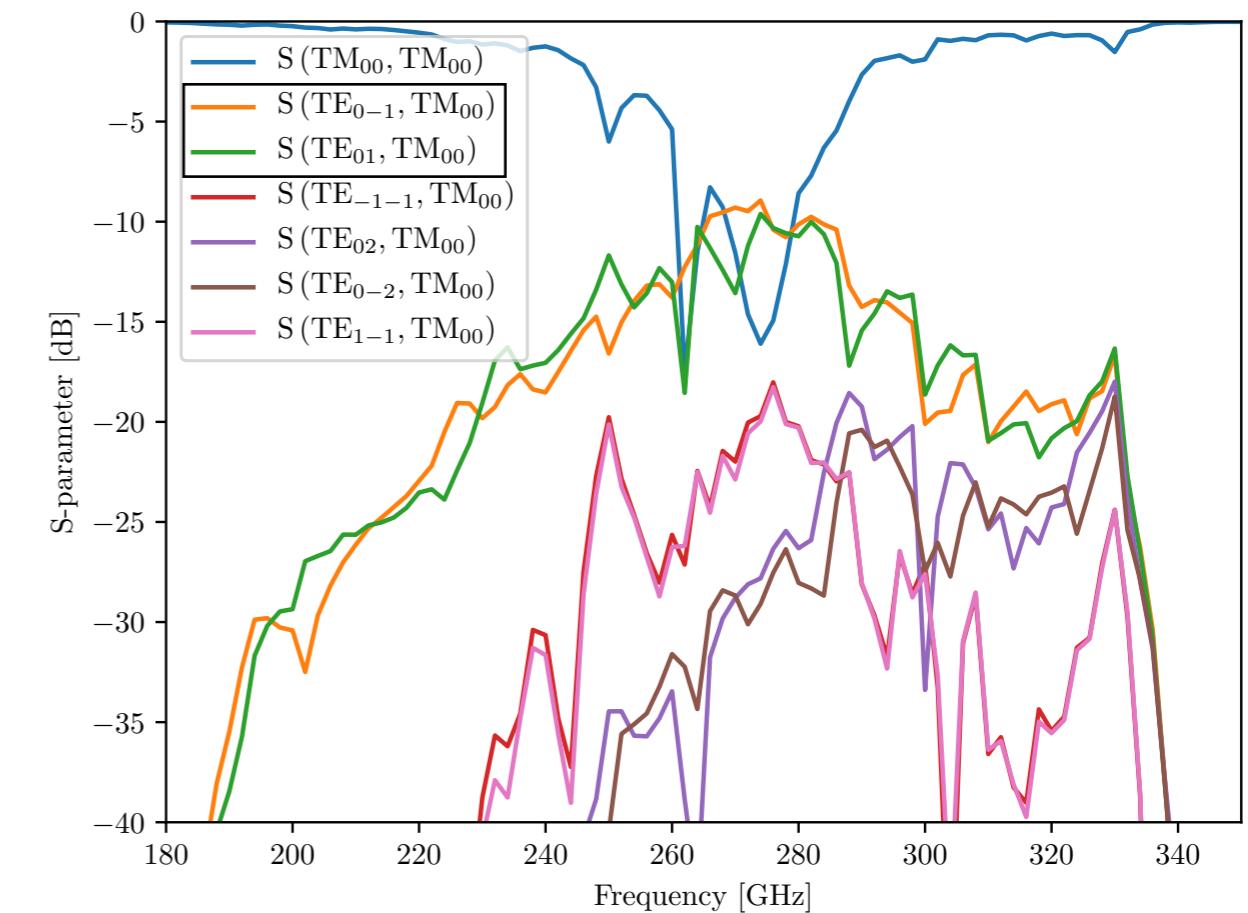


S-parameter

Pol 1



Pol 2



Modes with S-para > -20 dB are shown

**Max. 12.1% absorbed by TE10 and TE-10
230-290 GHz averaged: 4.3%**

**Max. 23.7% absorbed by TE10 and TE-10
230-290 GHz averaged: 10.1%**

Surface loss analysis

$$p_s = \operatorname{Re}(\vec{P} \cdot \vec{n})$$

P: Poynting vector

n: normal vector to the surface S

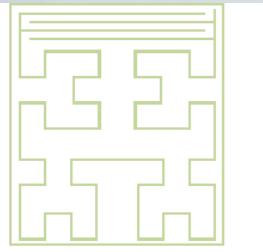
p_s: surface loss density

$$\text{Loss} = \int_S p_s dS$$

Surface loss could be estimated by integrating p_s with surface area

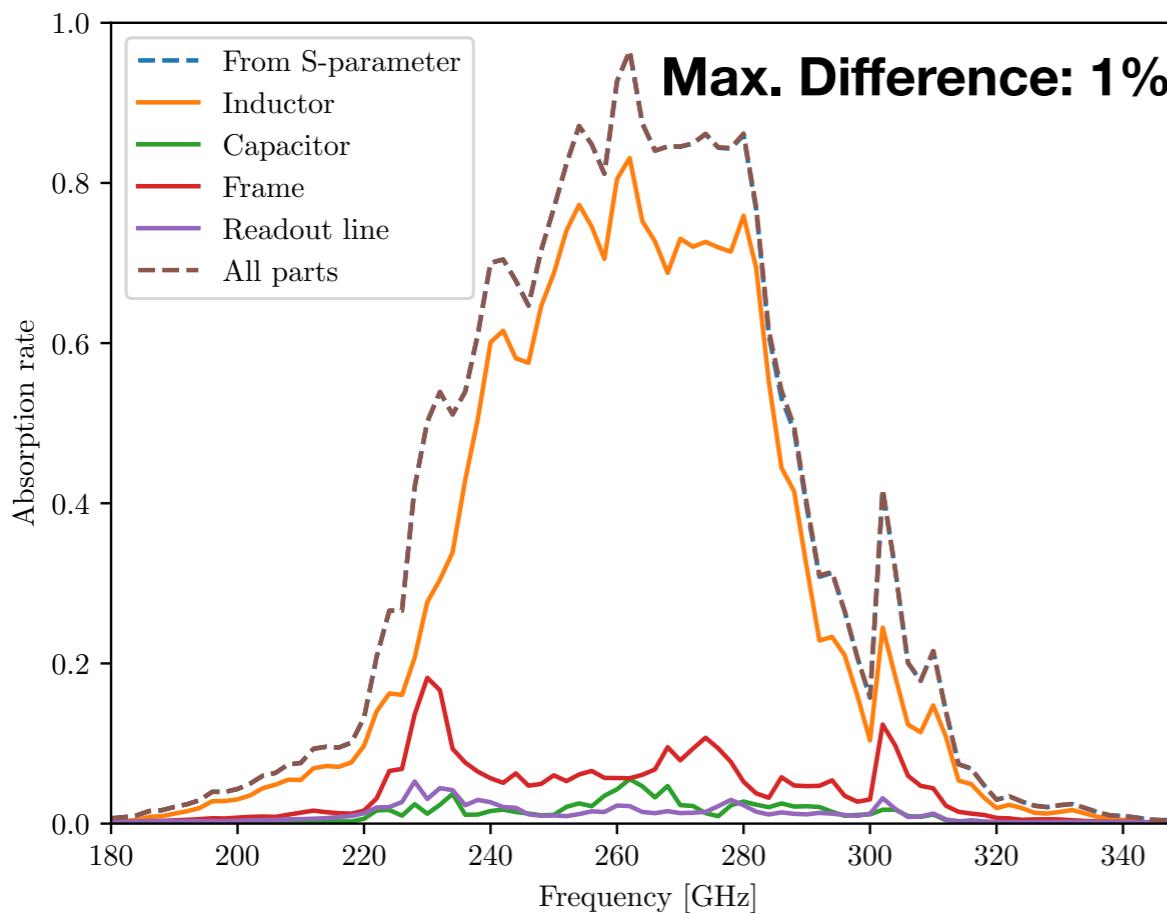
$$\text{Absorption rate} = \frac{\text{Loss}}{\text{Incident power}}$$

This calculation is done with “Field Calculator” in HFSS

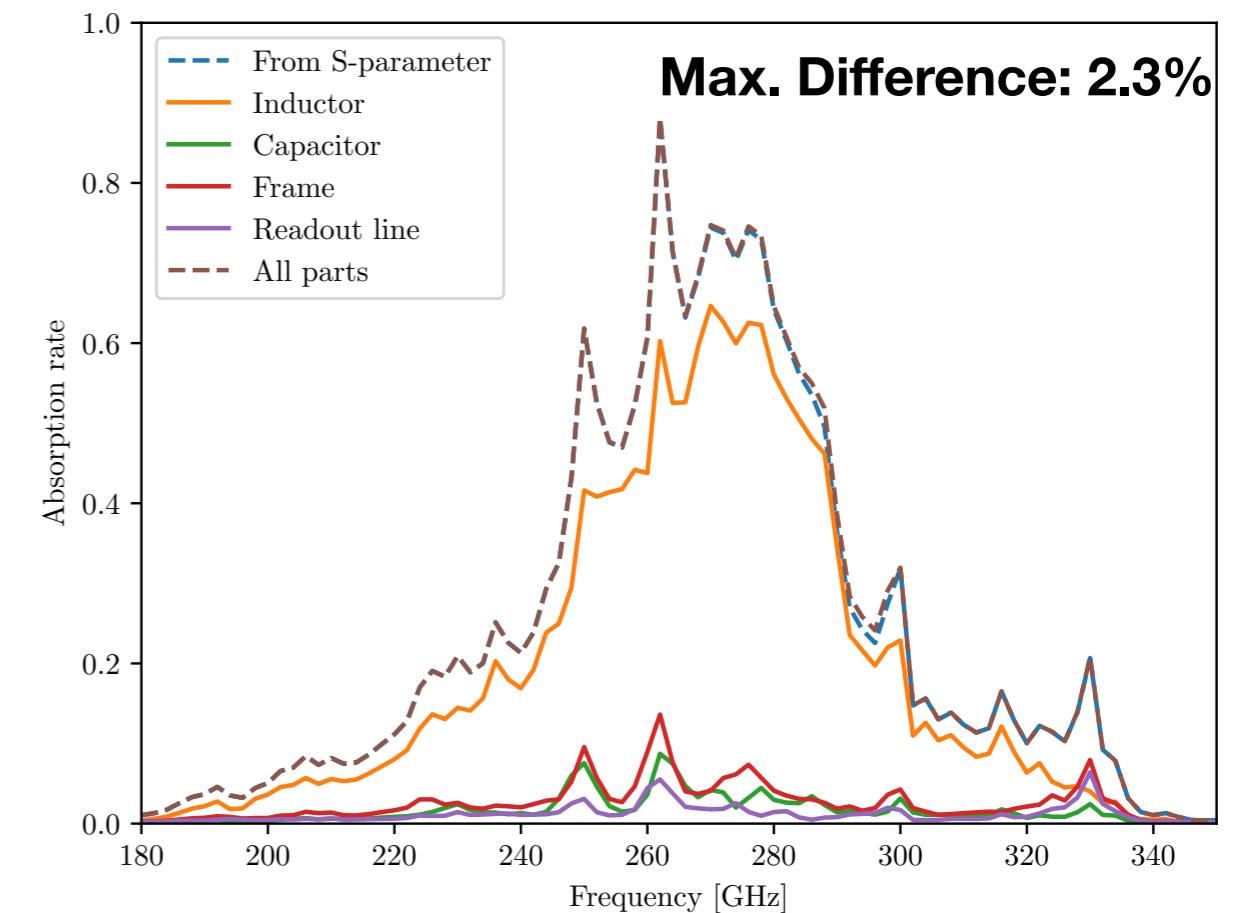


Surface loss analysis

Pol 1



Pol 2



230-290 GHz averaged: 61.7%

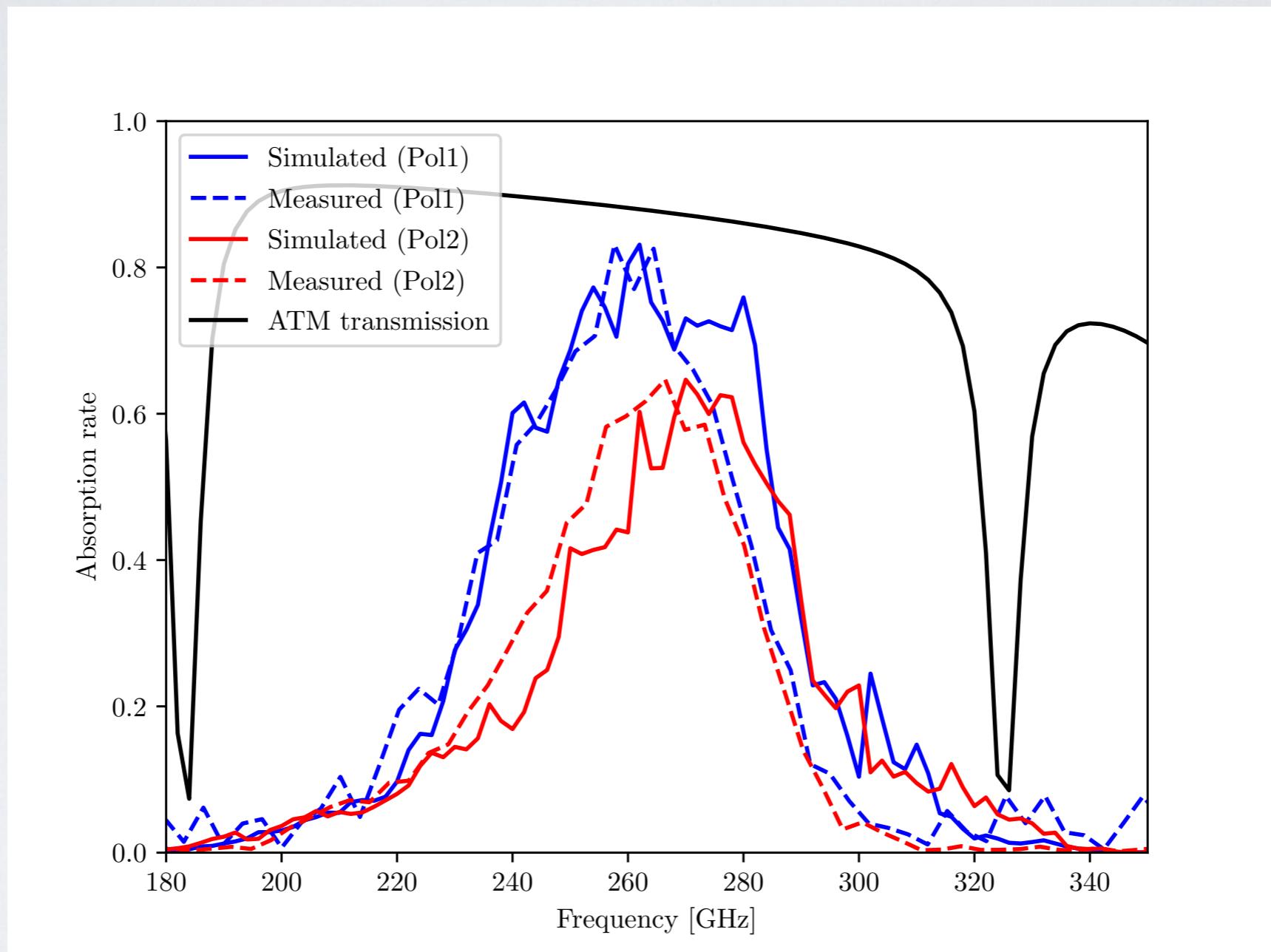
230-290 GHz averaged: 41.2%

Absorption from S-para:

$$1 - \sum_{n=1}^{all} S1(n)1(1)$$

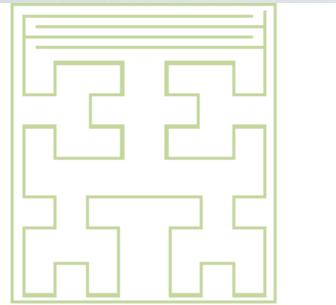
$$1 - \sum_{n=1}^{all} S1(n)1(2)$$

Compared with FTS results

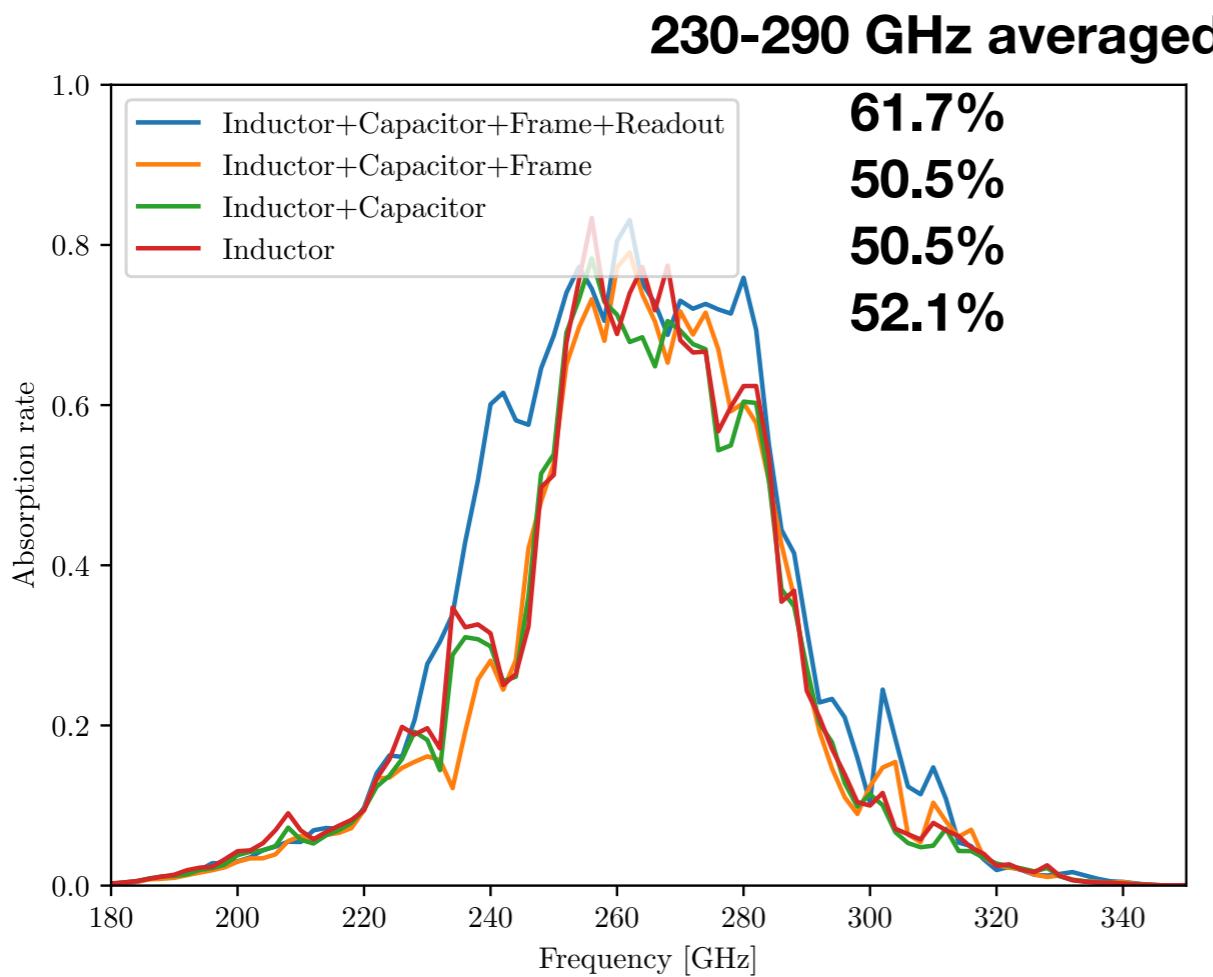


We assume that only absorption in inductor is detected

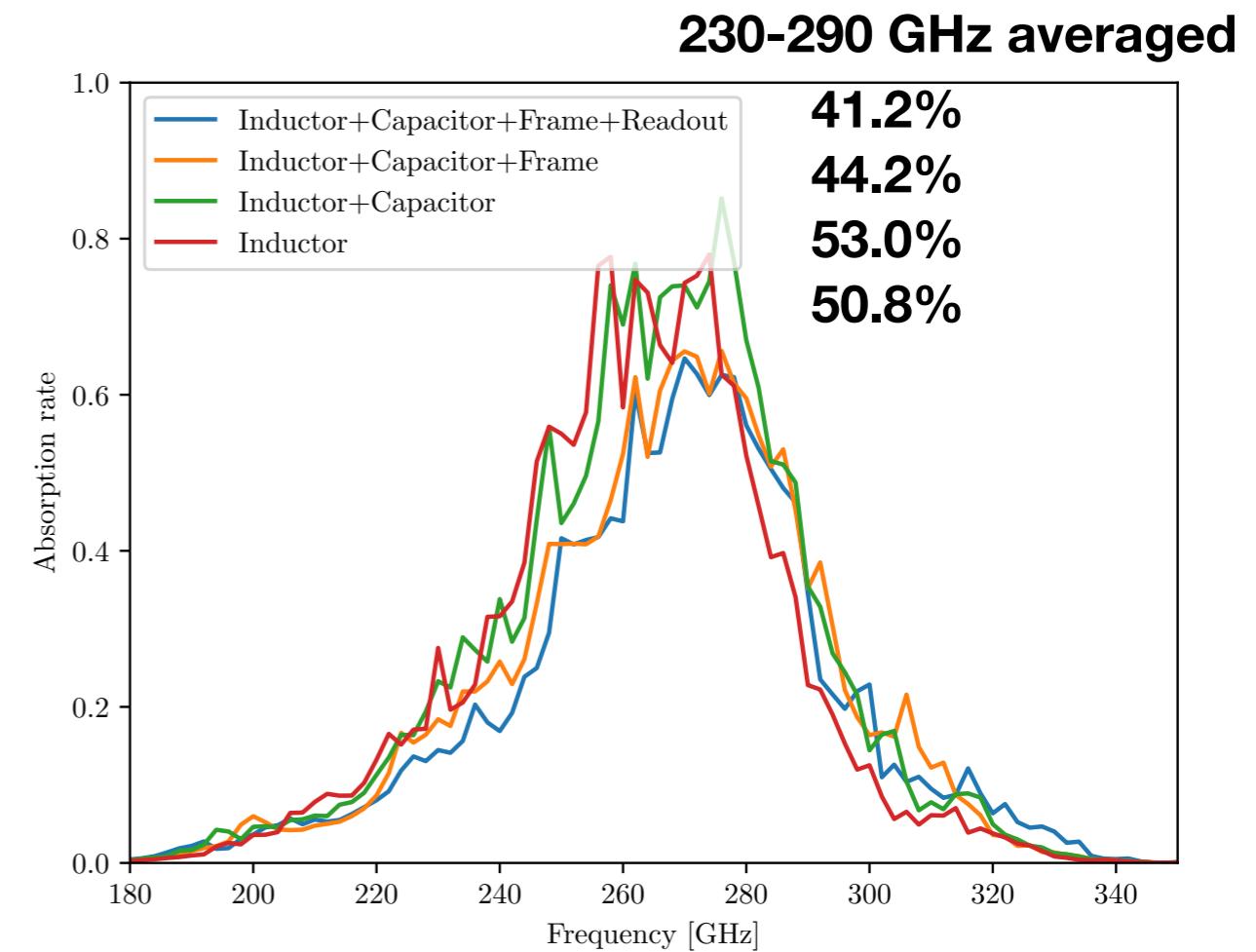
Effect of LEKID structure



Pol 1



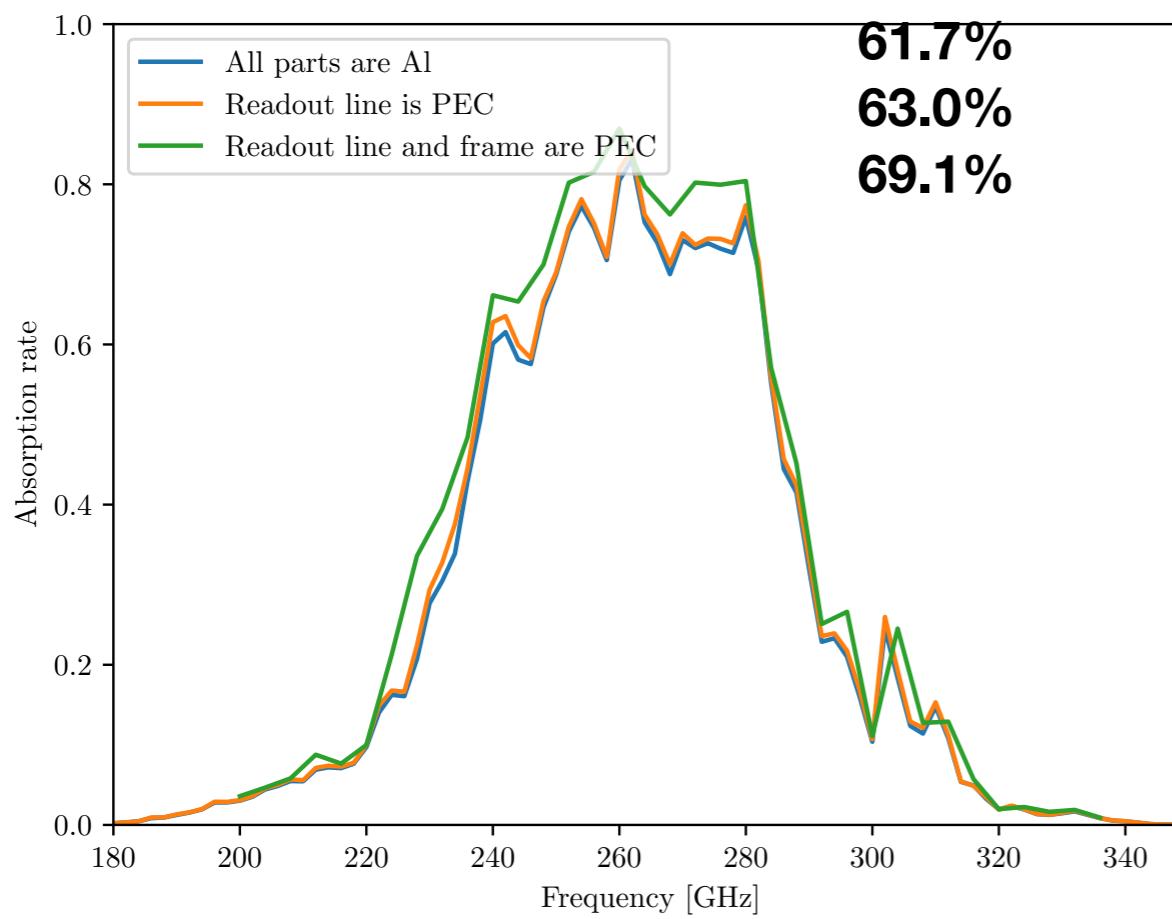
Pol 2



Could PEC (e.g. Nb) help?

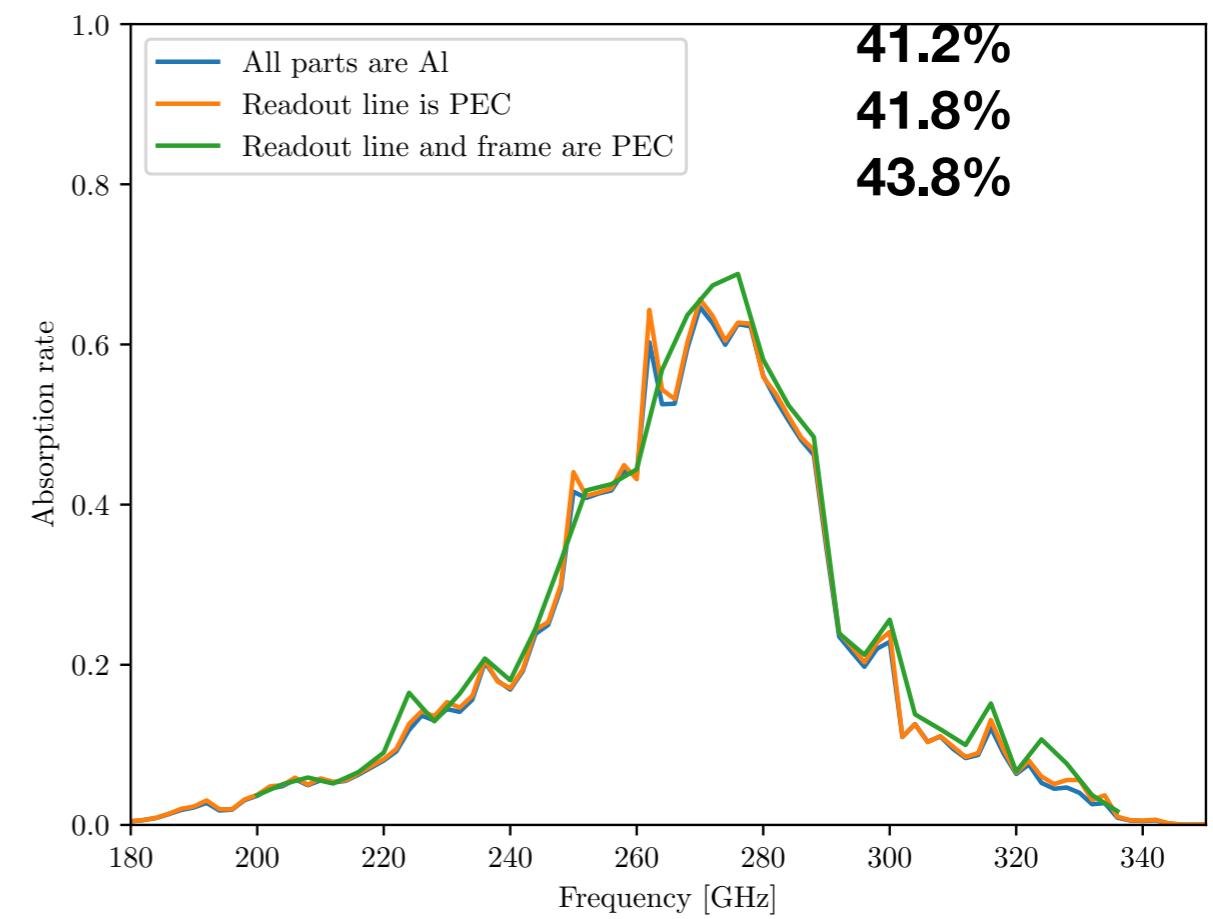
Pol 1

230-290 GHz averaged



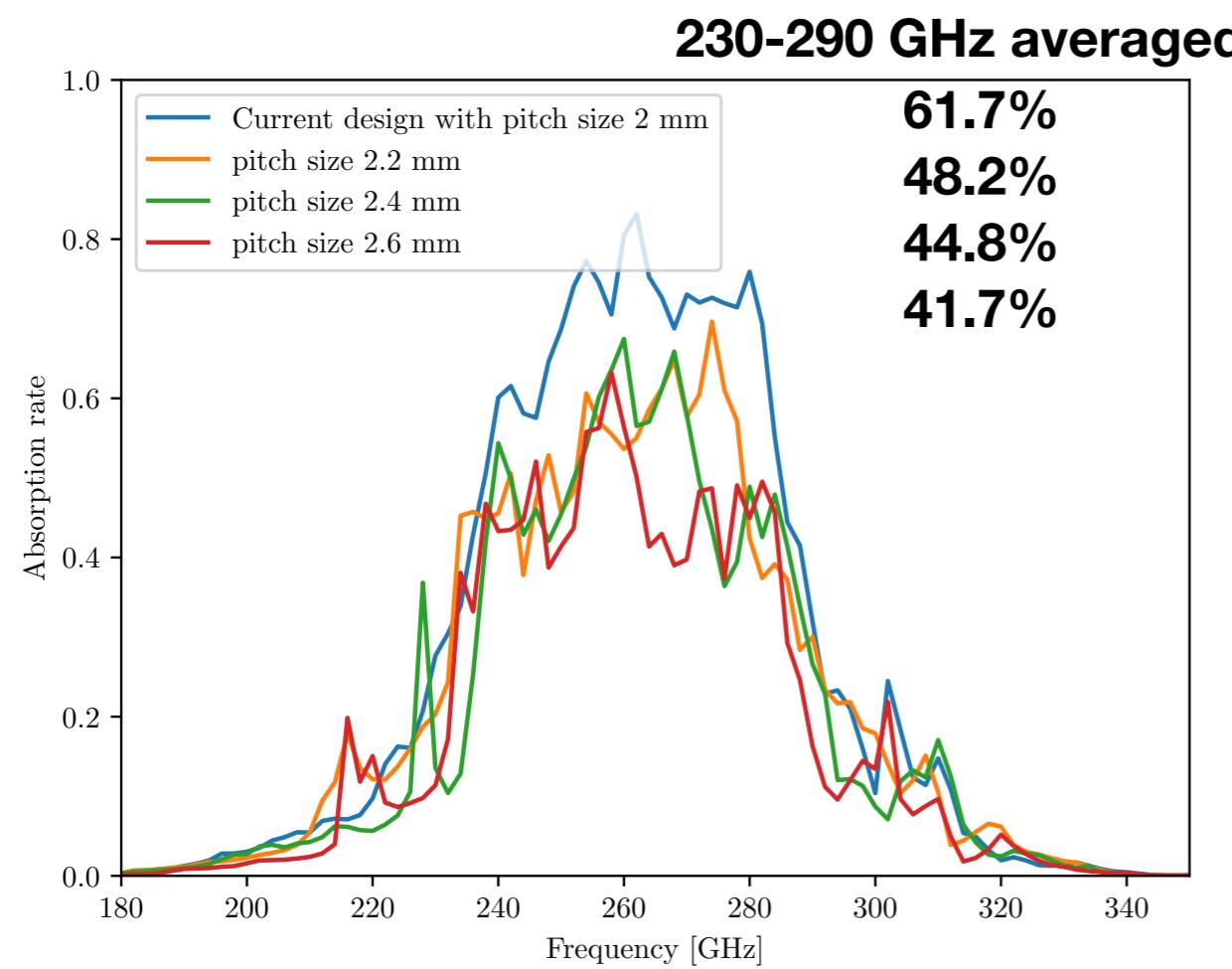
Pol 2

230-290 GHz averaged

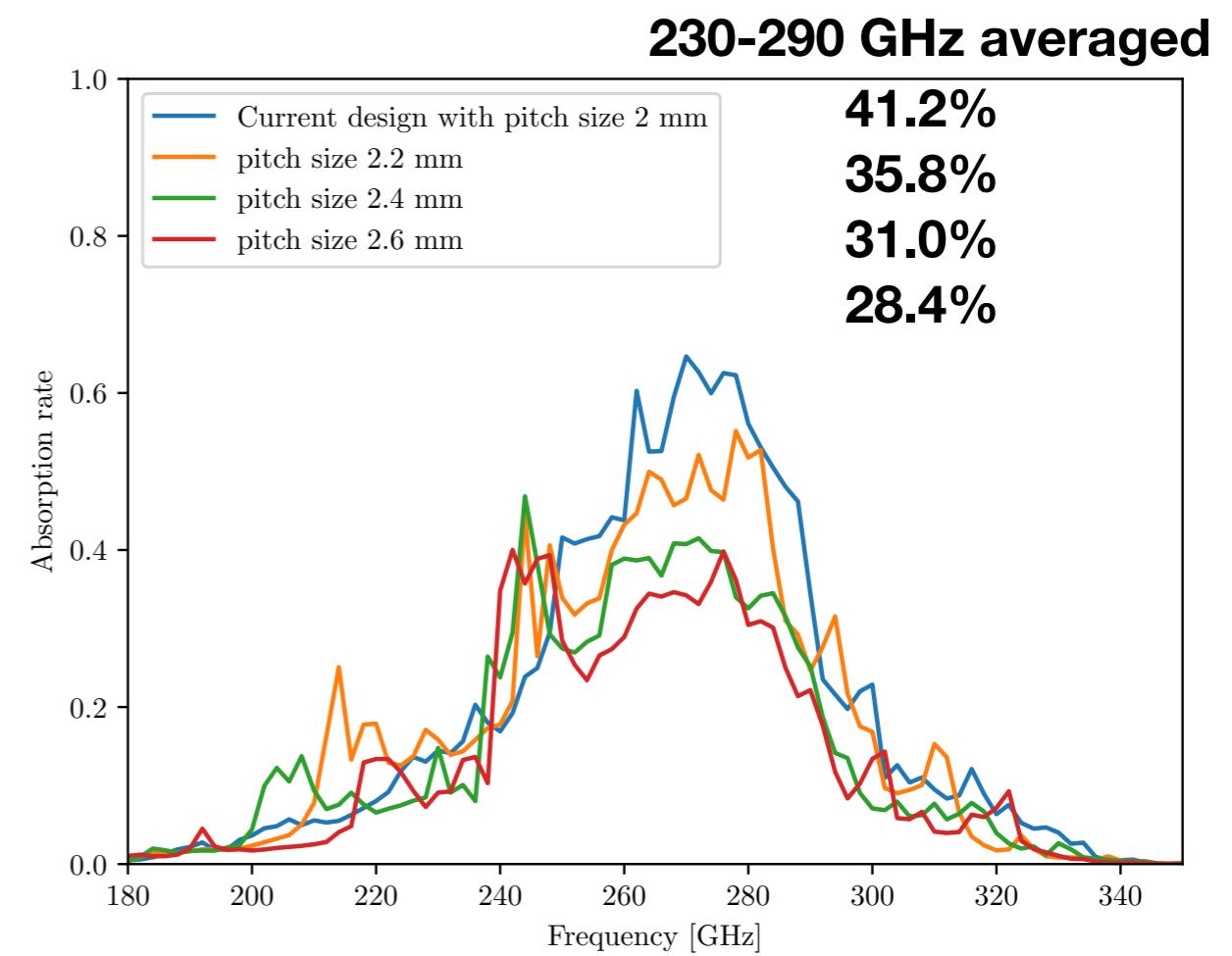


Pitch size matters?

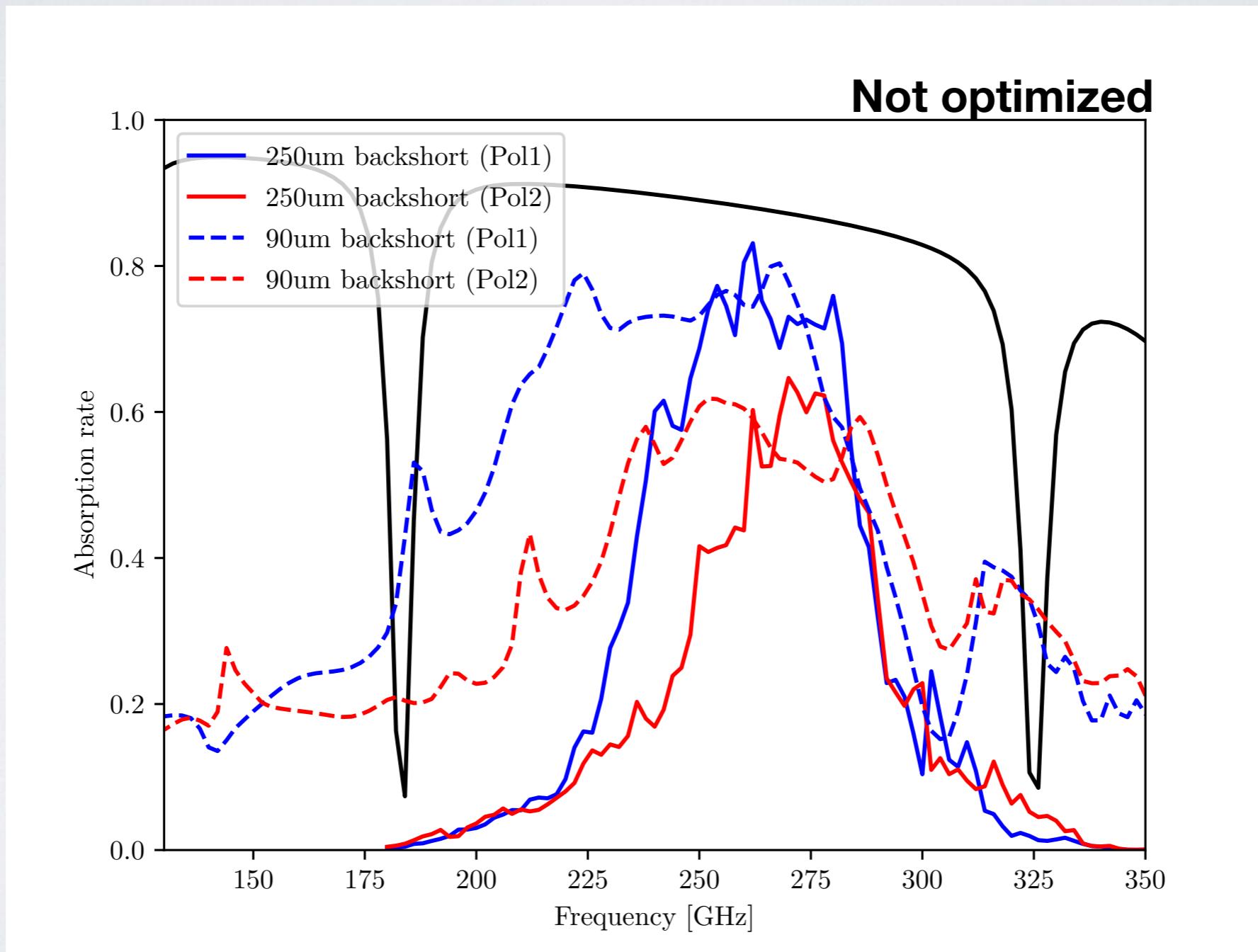
Pol 1



Pol 2



For larger band coverage



Conclusion

- A prototype high angular resolution LEKIDs array is designed and measured for NIKA2 1mm
Updated results: <https://arxiv.org/abs/1710.08232>
- The focal plane layout is pushed to a compact 1.4×1.4 mm² pixel size design
- A simulation method is demonstrated for LEKIDs array optical analysis to estimate energy absorbed in inductor
- Be careful with bi-polarization structure